



AEGIS

Collaboration

Meeting

Toledo,

9-11 December

2009

Reconciling a significant buildup of massive early-type galaxies through major mergers at $z \leq 1$ with mass-downsizing

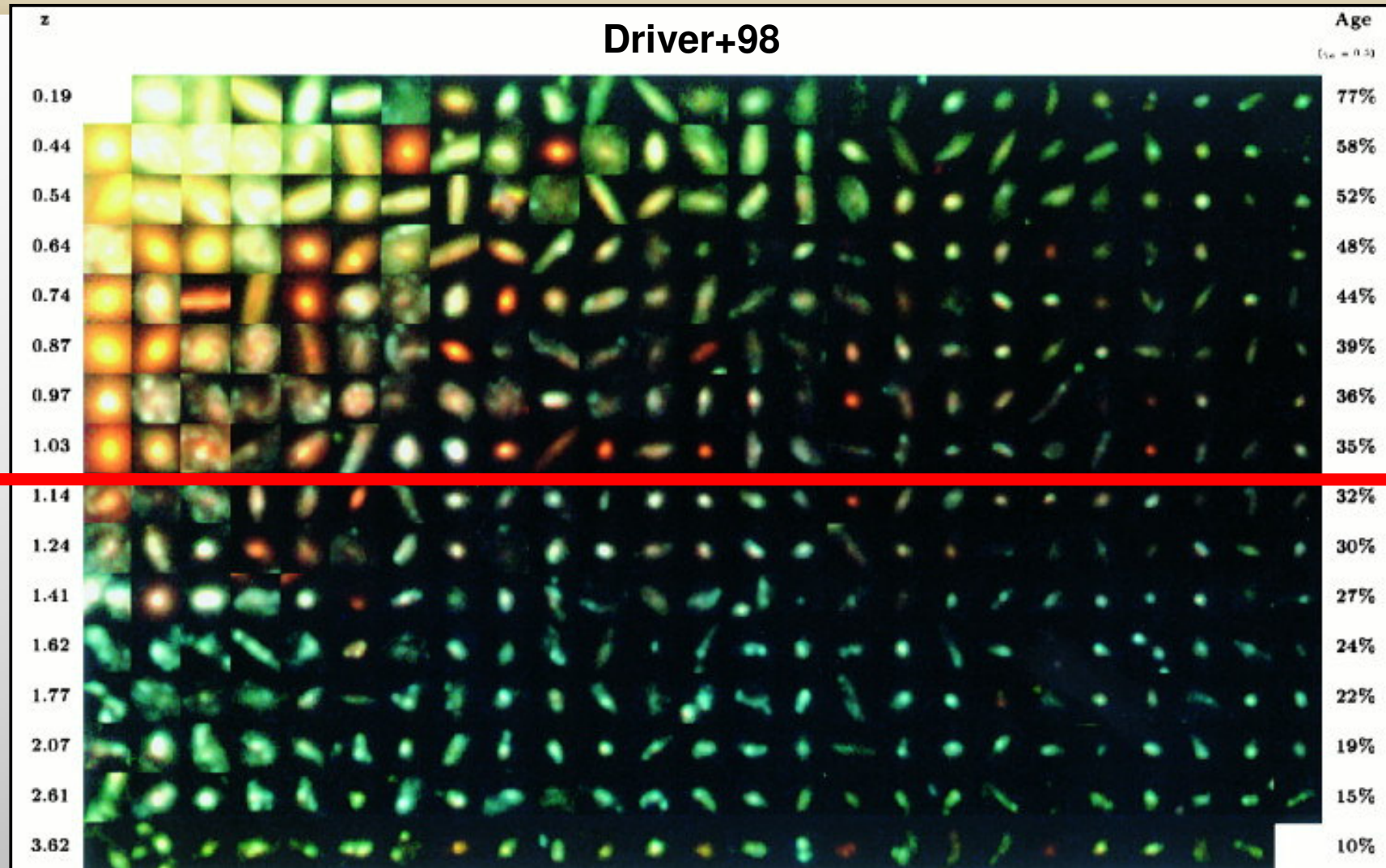
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(UCM), C. López-Sanjuan & M. Balcells (IAC), R.
Guzmán (UF), & J.C. Muñoz-Mateos (UCM)



I- Introduction

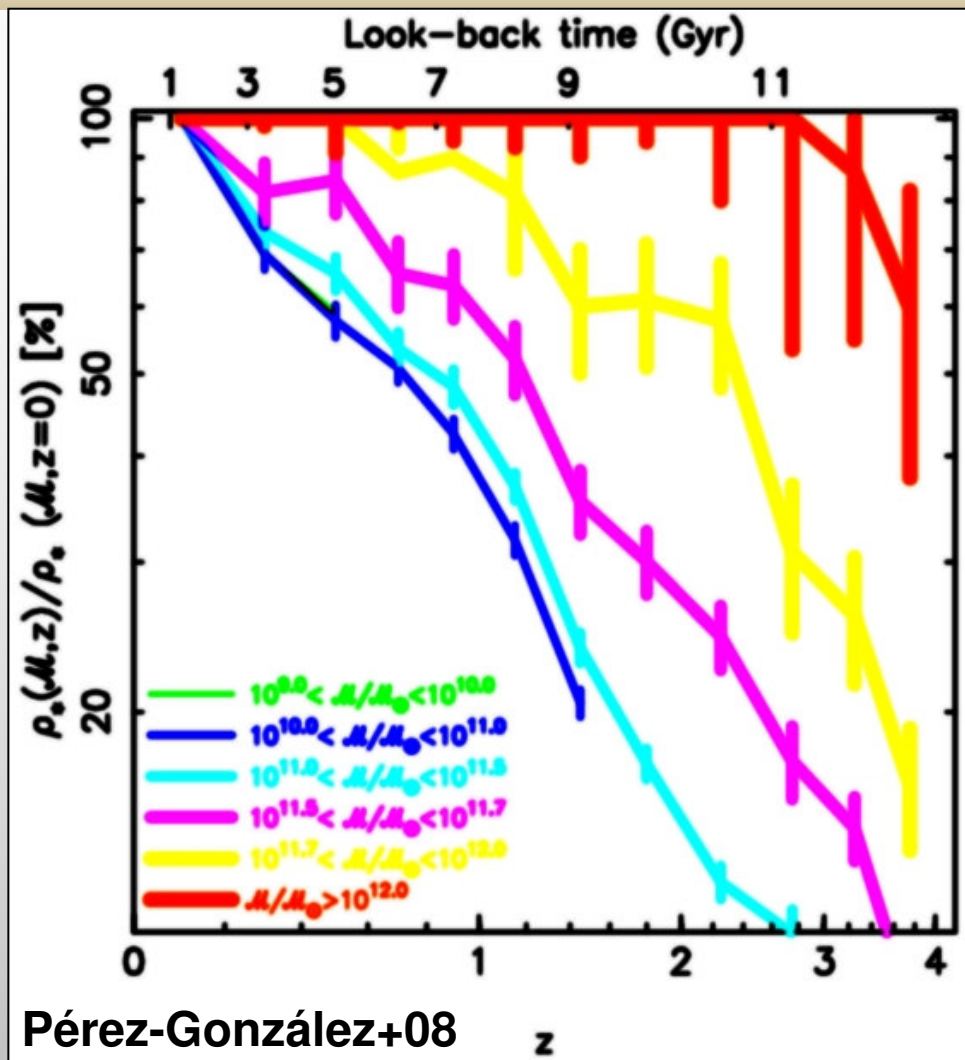
Assembly of early-type galaxies





I- Introduction

Galaxy mass-downsizing



- ✓ Downsizing in mass and the hierarchical scenario, in apparent contradiction.
- ✓ We propose a direct test to the hierarchical origin of massive early-type galaxies (ETGs):

How would have evolved the population of massive ETGs backwards in time, assuming that they derive from the major mergers that are strictly reported by observations?

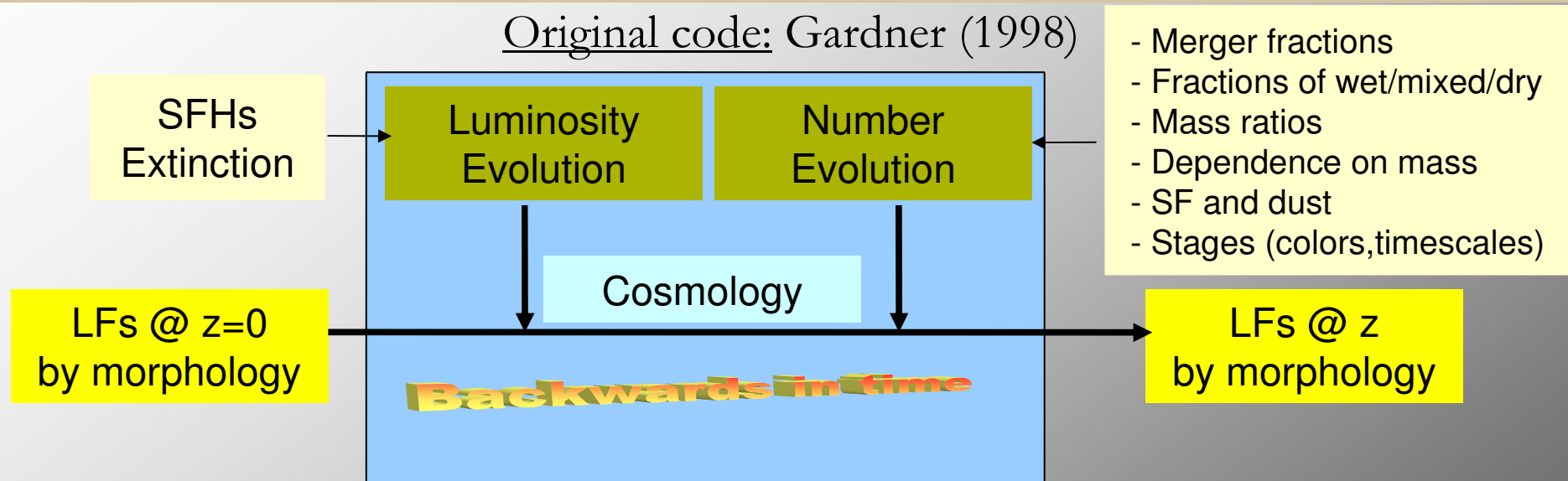


II- Models

General overview



Original code: Gardner (1998)



Main hypotheses:

- ✓ Each major merger gives place to an ETG.
- ✓ Gas-rich major mergers undergo phases of dust-reddened, star-forming galaxies (DSFs).

Note: Limiting magnitude due to observational merger fractions:

$M(B) < -20$ mag (\Rightarrow ETGs with $\log(\text{Mass}/M_{\text{sun}}) > 10.7$, Cimatti et al. 2006).

Eliche-Moral et al., ApJ submitted



II- Models

Merging procedure



Wet merger

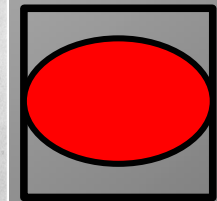
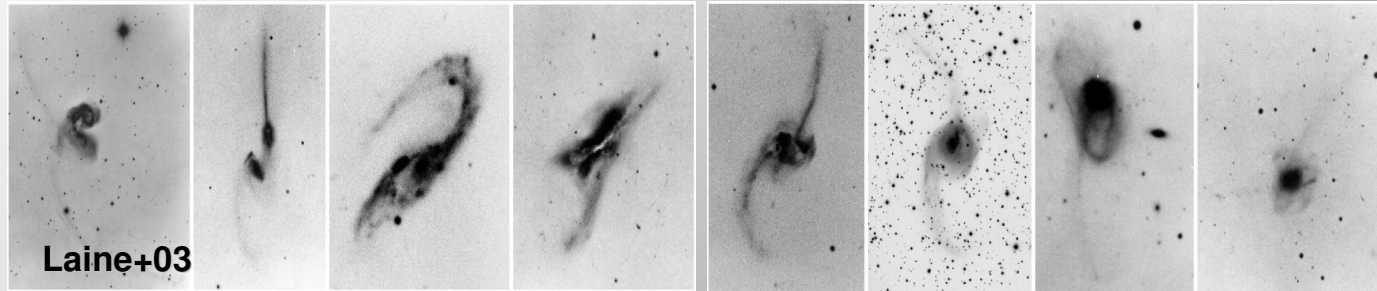
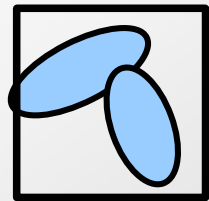
time

2 Sc-Irr's

2 DSFs

1 DSF

1ETG



Mixed merger

time

1 Sc-Irr

1 DSF

1 DSF

1ETG

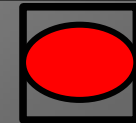


1ETG

1ETG

1 DSF

1ETG



Dry merger

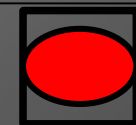
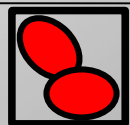
time

2ETGs

2ETGs

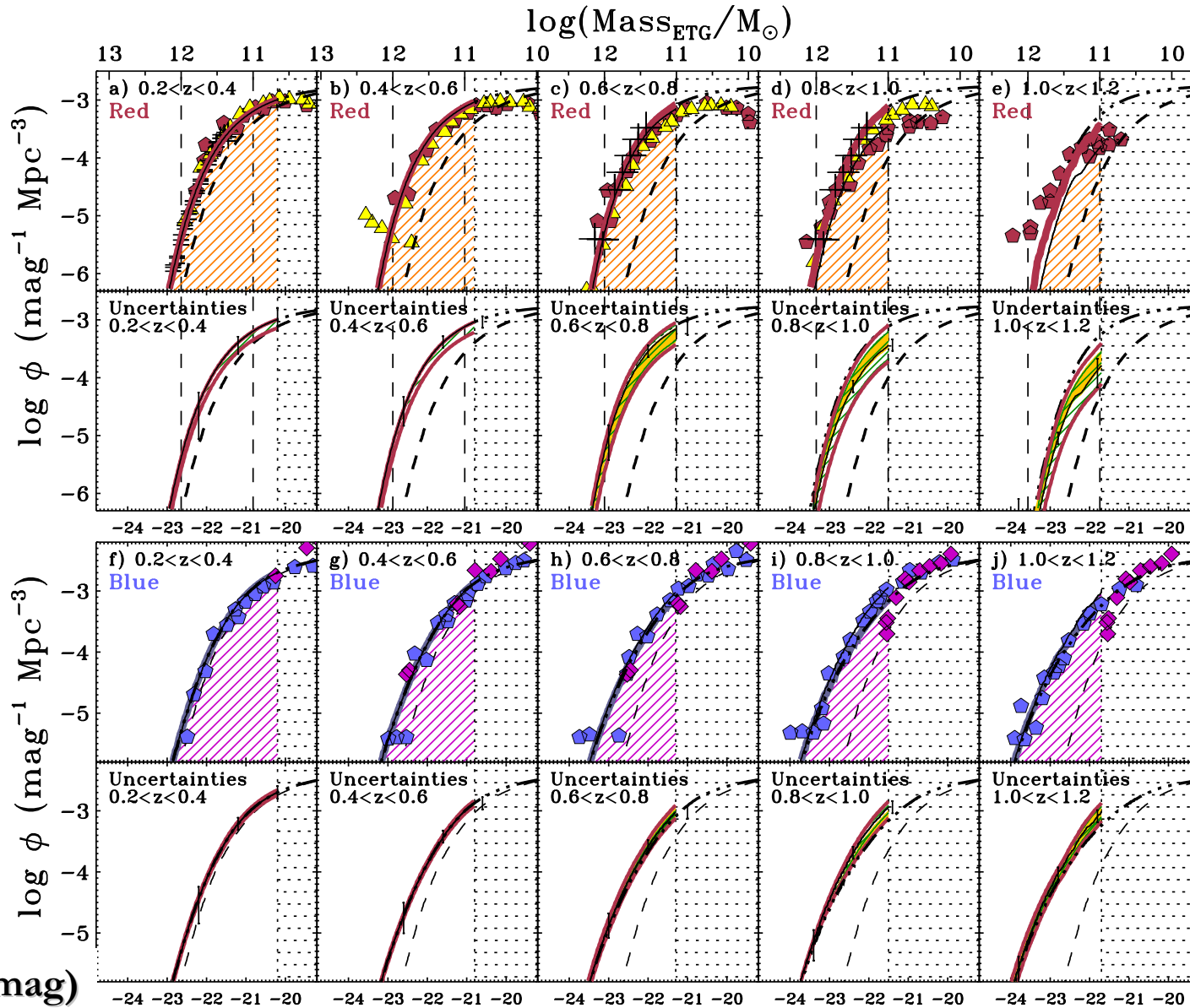
1ETG

1ETG





III- Results: B-band LFs by color up to $z \sim 1$



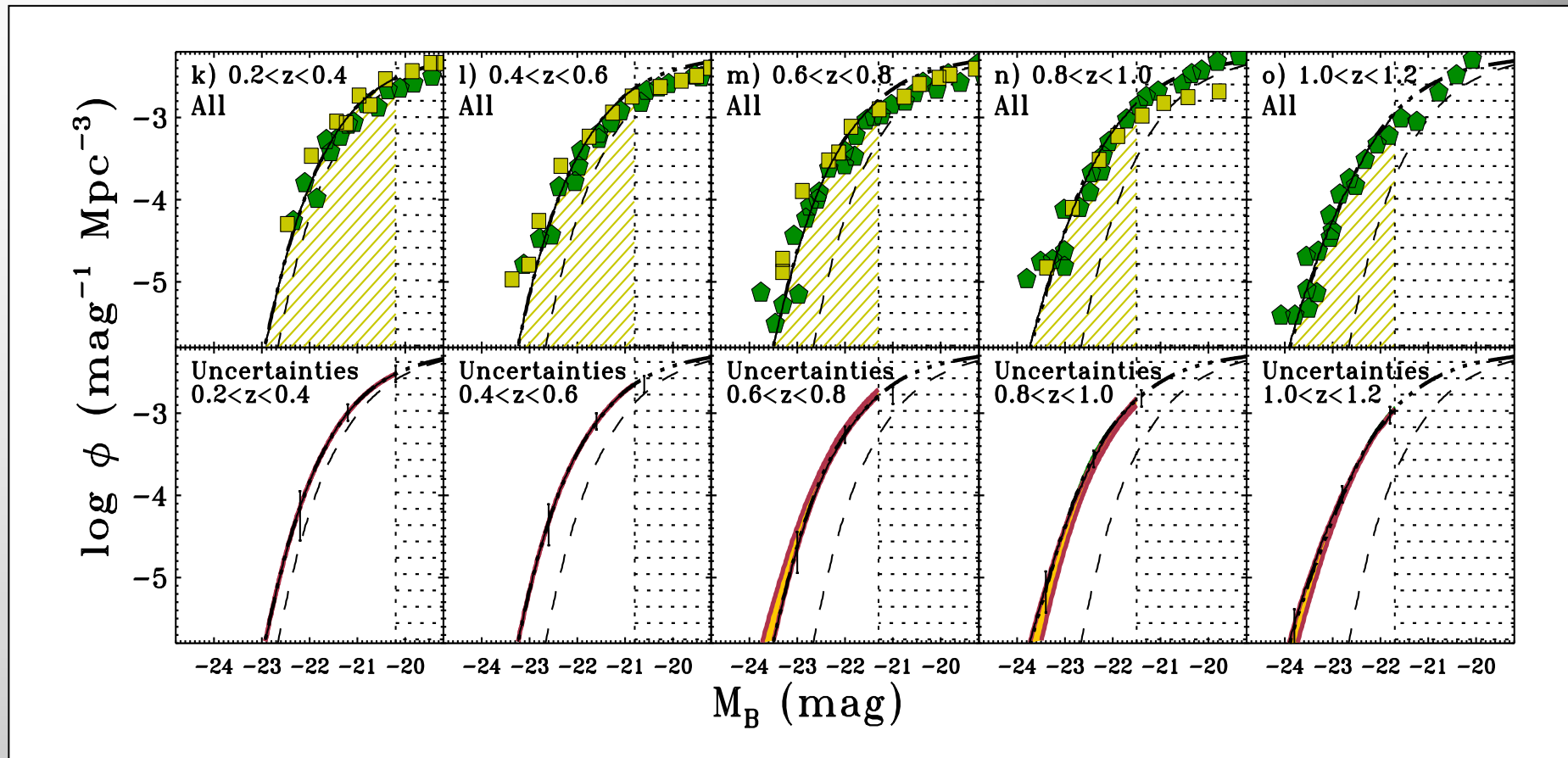
M(B) (mag)



III- Results:

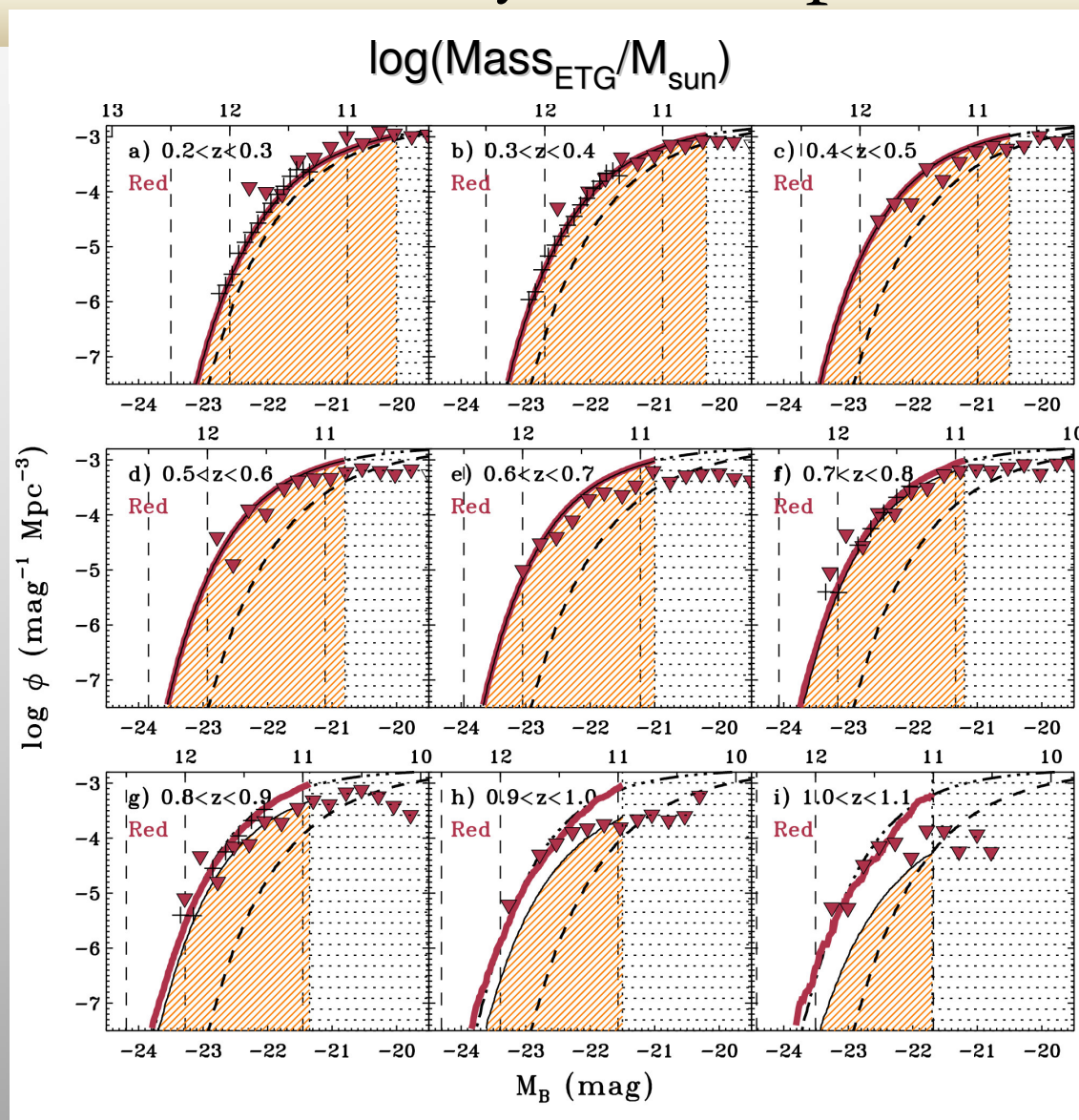


B-band LFs by color up to $z \sim 1$



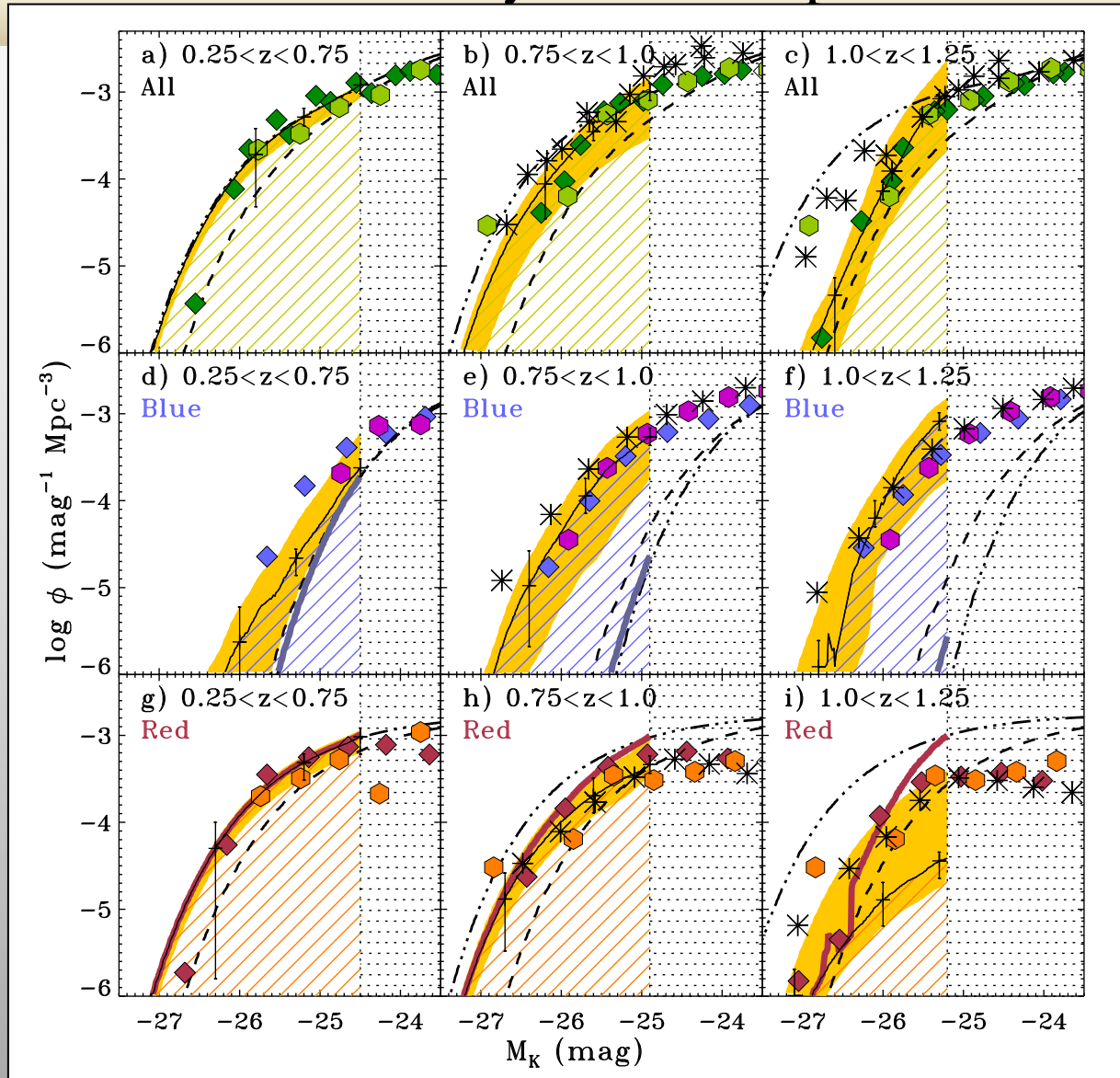


III- Results: B-band LFs by color up to $z \sim 1$





III- Results: K-band LFs by color up to $z \sim 1$

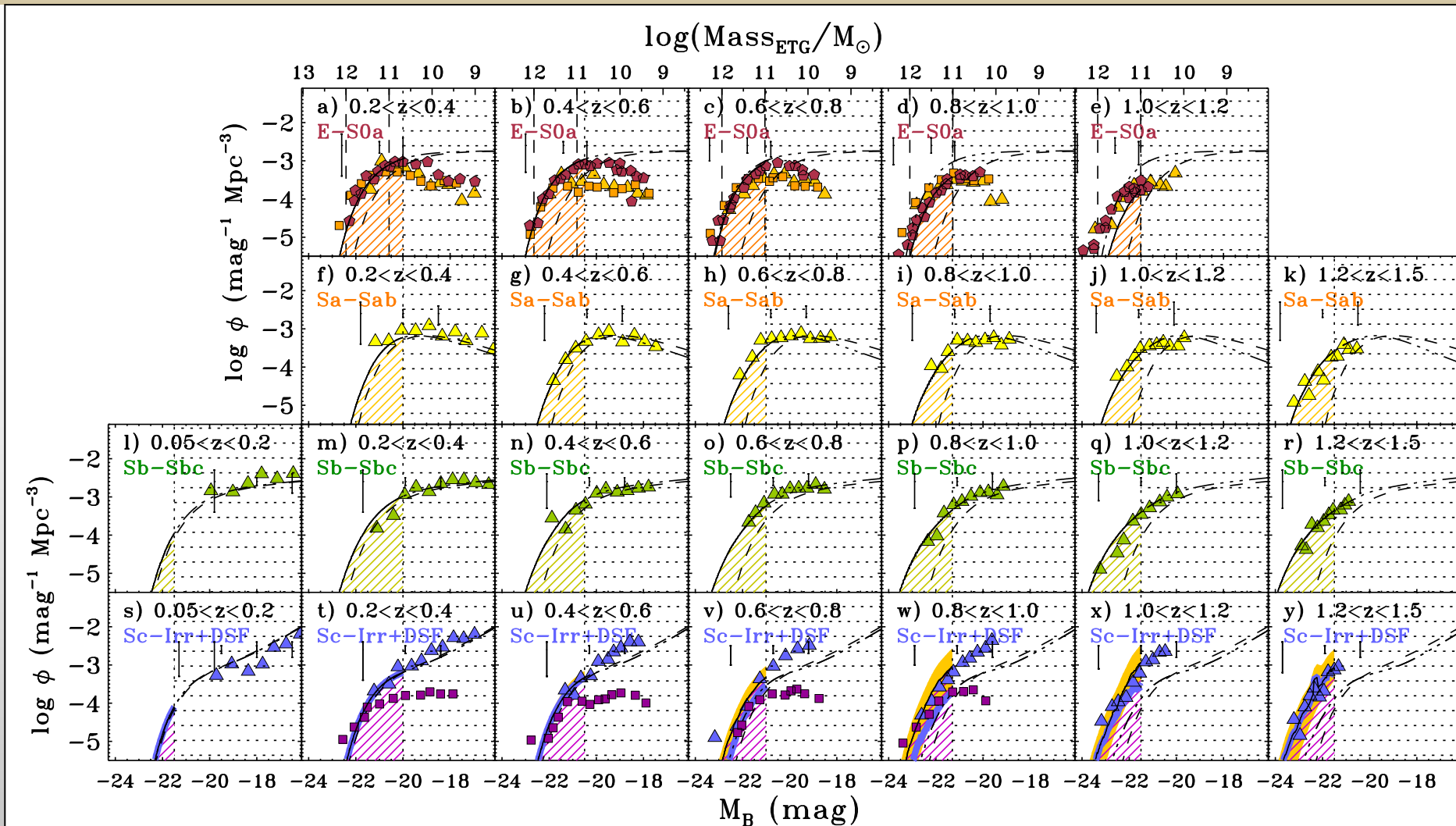




III- Results:



B-band LFs by morphology up to $z \sim 1$





IV- Conclusions



- 1) The model **predicts the evolution of the observational LFs up to $z \sim 1$** simultaneously in different bands, for different selection criteria on color and morphology.
- 2) It is feasible to **build up $\sim 50-60\%$ of present-day massive ETGs since $z \sim 1$ through major mergers strictly reported by observations.** The predicted gas-rich progenitors explain naturally the blue galaxy excess at $z > 0.8$.
- 3) It predicts observational **trends in agreement with mass-downsizing at $z < 1$,** considering hierarchical assembly of ETGs. It provides a framework that reconciles different observational results on the LFs of red galaxies.
- 4) It also predicts that **$z \sim 0.8$ is a transition epoch for this buildup,** and the appearance of a significant population of **DSFs at $z > 0.8$,** in agreement with observations.

Major mergers must have been the main driver for the observed migration of mass from the blue galaxy cloud to the red sequence since $z \sim 1$,
for galaxy masses $> 5 \times 10^{10} M_{\text{sun}}$.